### A&OS 101 - Accelerations owing to sphericity

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Before, we considered coordinate system rotation and divided the acceleration as seen from space  $(\frac{d_a\vec{U}_a}{dt})$  into these parts: the acceleration seen from the Earth  $(\frac{d\vec{U}}{dt})$  and terms relating to Earth rotation (Coriolis and centrifugal accelerations). Now we have a rotation of coordinates due to the sphericity of the Earth. This will entail splitting  $\frac{d\vec{U}}{dt}$  into two parts, representing accelerations relative to a flat Earth and compensation for Earth curvature.

#### Velocity components

Our unit vectors  $\hat{i}$ ,  $\hat{j}$ , and  $\hat{k}$  remain pointing east, north and up, but we need to move from Cartesian position x, y and z to spherical position  $\lambda$ ,  $\phi$  and z where  $\lambda$  is longitude and  $\phi$  is latitude. Thus, we need to relate velocities u and v in terms of  $\frac{d\lambda}{dt}$  and  $\frac{d\phi}{dt}$ .



Figure 1: Meridional velocity v as a tangential velocity.

Figure 1 shows that the north-south (meridional) velocity v along a given longitude  $\lambda$  can be interpreted as a velocity tangent to a circle of distance r from the center of the Earth. (This r = a + z, where a is the Earth's radius and z is distance above the Earth's surface.) Recall that tangential velocity is is angular velocity times radius. The angular velocity is  $\frac{d\phi}{dt}$  since latitude is changing with time. Therefore, we have

$$v = r \frac{d\phi}{dt}.$$
(1)

Now look at east-west (zonal) velocity u, representing a longitude change  $d\lambda$  with time along a latitude circle (Fig. 2). It is clear that u is also a tangential velocity. The radius of the latitude circle is R, distance from the spin axis, which may also be expressed as  $r \cos \phi$ . It follows that

$$u = R \frac{d\lambda}{dt} = r \cos \phi \frac{d\lambda}{dt}.$$
(2)

Finally,  $w = \frac{dz}{dt}$  on both flat and spherical Earths.



Figure 2: Zonal velocity u as a tangential velocity.

### Chain rule

With regard to 3D vector velocity,  $\vec{U} = u\hat{i} + v\hat{j} + w\hat{k}$ , six aspects of this expression could change with time: the component velocities, and the coordinate axes. Thus the chain rule applied to  $\frac{d\vec{U}}{dt}$  yields

$$\frac{d\vec{U}}{dt} = \frac{du}{dt}\hat{i} + \frac{dv}{dt}\hat{j} + \frac{dw}{dt}\hat{k} + u\frac{d\hat{i}}{dt} + v\frac{d\hat{j}}{dt} + w\frac{d\hat{k}}{dt}.$$
(3)

We need to find expressions for the coordinate axis accelerations.

This effort starts with identifying the dimensions in which each axis varies. The  $\hat{i}$  axis is simplest since it is a function only of longitude. Moving along any given latitude circle,  $\hat{i}$  changes direction owing to Earth's curvature, as shown in Fig. 3. The other two coordinate axes vary in both latitude and longitude, as shown in Figures 4 and 5.



Figure 3: The  $\hat{i}$  coordinate axis varies with longitude only.



Figure 4: The  $\hat{j}$  coordinate axis varies with longitude and latitude.



Figure 5: The  $\hat{k}$  coordinate axis varies with longitude and latitude.

## The $\hat{i}$ axis

The  $\hat{i}$  unit vector always points east, but what we define as "east" changes around a latitude circle. Indeed,  $\hat{i}$  can be interpreted as tangential velocity and its change  $\Delta \hat{i}$  as a centripetal acceleration, as shown in Fig. 6. Recognizing that  $\hat{i}$  varies only longitudinally, the chain rule applied to  $\frac{d\hat{i}}{dt}$  quickly reduces in the following manner:

$$\frac{d\hat{i}}{dt} = \frac{\partial\hat{i}}{\partial t} + u\frac{\partial\hat{i}}{\partial x} + v\frac{\partial\hat{i}}{\partial y} + w\frac{\partial\hat{i}}{\partial z}$$
(4)

$$= u \frac{\partial \hat{i}}{\partial x}.$$
 (5)

We will need  $\frac{d\hat{i}}{dt}$ , and that is

$$d\frac{d\hat{i}}{dt} = u^2 \frac{\partial\hat{i}}{\partial x}.$$
(6)

The derivation continues to mimic that done for centripetal acceleration. The magnitude of  $\frac{\partial \hat{i}}{\partial x}$ is  $\frac{\Delta \hat{i}}{\Delta x}$ , where dx is the arclength over angle  $d\lambda$  (i.e.,  $dx = r \cos \phi \Delta \lambda$ ). For small  $d\lambda$ ,  $\Delta \hat{i} \approx d\lambda$  (recall unit vectors have unit length). Therefore,

$$\frac{\Delta \hat{i}}{\Delta x} = \frac{\Delta \lambda}{r \cos \phi \Delta \lambda} = \frac{1}{r \cos \phi}.$$
(7)

Just as with centripetal acceleration, we note  $\delta \hat{i}$  points inward towards the spin axis, and has two components – northward and towards Earth's center – as shown in Fig. 7. The northward



Figure 6: An augmented version of Fig. 3, showing the change of the  $\hat{i}$  axis along a latitude circle is interpretable as a centripetal acceleration.

component is  $\Delta \hat{i} \sin \phi \hat{j}$  and the other is  $-\Delta \hat{i} \cos \phi \hat{k}$ . As a result,

$$\frac{\partial \hat{i}}{\partial x} = \frac{1}{r\cos\phi} \left( \hat{j}\sin\phi - \hat{k}\cos\phi \right),\tag{8}$$

leading to

$$u^{2}\frac{\partial\hat{i}}{\partial x} = \frac{u^{2}}{r\cos\phi}\left(\hat{j}\sin\phi - \hat{k}\cos\phi\right).$$
(9)

This creates a  $\frac{u^2}{r} \tan \phi$  term for the v equation of motion and a  $\frac{u^2}{r}$  term for the w equation.



Figure 7:  $\Delta \hat{i}$  and its components.

# The $\hat{j}$ and $\hat{k}$ equations

The other two coordinates are more complicated, owing to dependence on both longitude and latitude, but the accelerations are derived in a very similar way. For the  $\hat{j}$  unit vector, we obtain

$$u\frac{\partial\hat{j}}{\partial x} = -\frac{u\tan\phi}{r}\hat{i},\tag{10}$$

$$v\frac{\partial \hat{j}}{\partial y} = -\frac{v}{r}\hat{k} \tag{11}$$

for the latitudinal and longitudinal dependences, respectively. The  $\hat{k}$  component results in

$$\frac{d\hat{k}}{dt} = \frac{u}{r}\hat{i} + \frac{v}{r}\hat{j}.$$
(12)

Therefore, (3) expands into

$$\frac{d\vec{U}}{dt} = \left[\frac{du}{dt} - \frac{uv\tan\phi}{r} + \frac{uw}{r}\right]\hat{i} \\
+ \left[\frac{dv}{dt} + \frac{u^2\tan\phi}{r} + \frac{vw}{r}\right]\hat{j} \\
+ \left[\frac{dw}{dt} - \frac{u^2 + v^2}{r}\right]\hat{k}.$$