

# Ion cyclotron waves at the Moon and their connection to the plasma sheet and the lunar exosphere

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1. UCLA; 2. Also at NASA/GSFC

Acknowledgments:

Bill Farrell, Jasper Halekas, Chris Russell, Hanying Wei

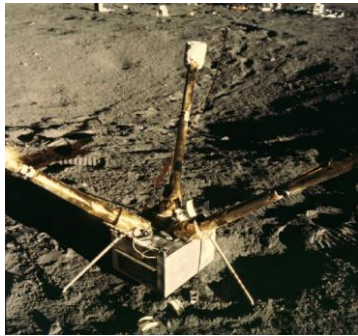
2015 GEM Summer Workshop

Tail Environment and Dynamics at Lunar Distances Focus Group

Snowmass, Colorado, June 14-19

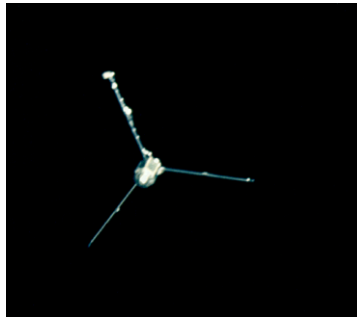
# Apollo Magnetic Field Experiments

## Lunar Surface Magnetometer (LSM)



Apollo 12  
 Apollo 15  
 Apollo 16  
 (NASA Ames)

## Sub-satellite Biaxial Magnetometer (SBM)



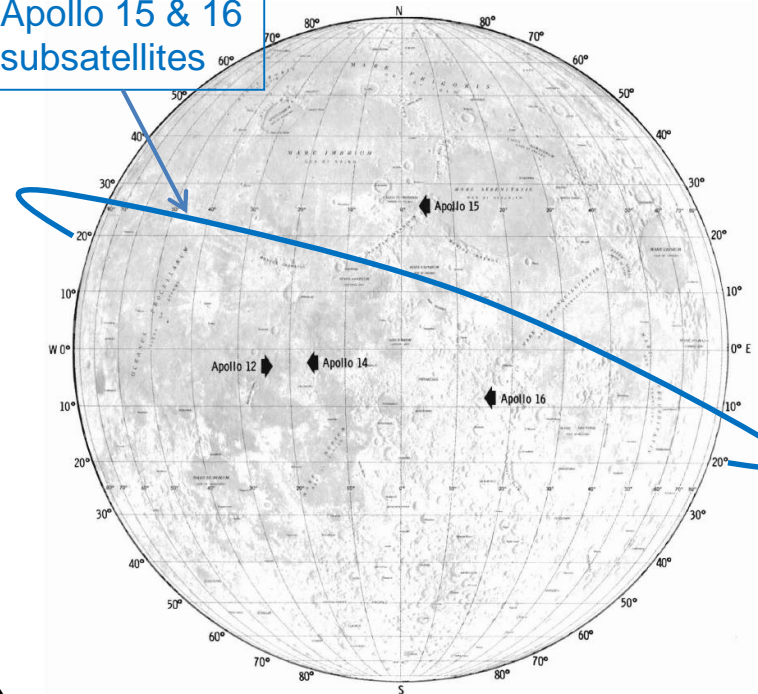
Apollo 15  
 Apollo 16  
 (UCLA)

## Lunar Portable Magnetometer

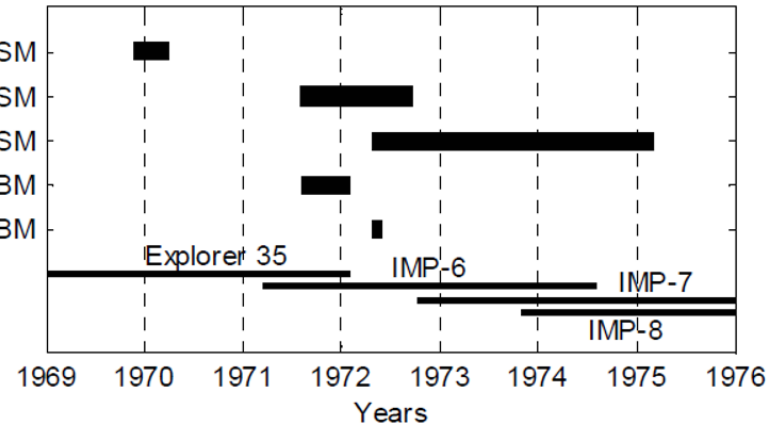


Apollo 14  
 Apollo 16  
 (NASA Ames)

Apollo 15 & 16 subsatellites



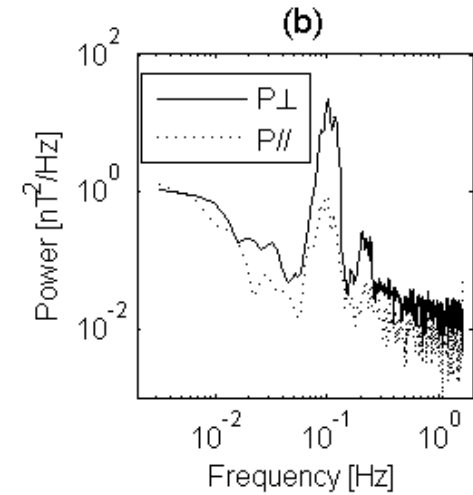
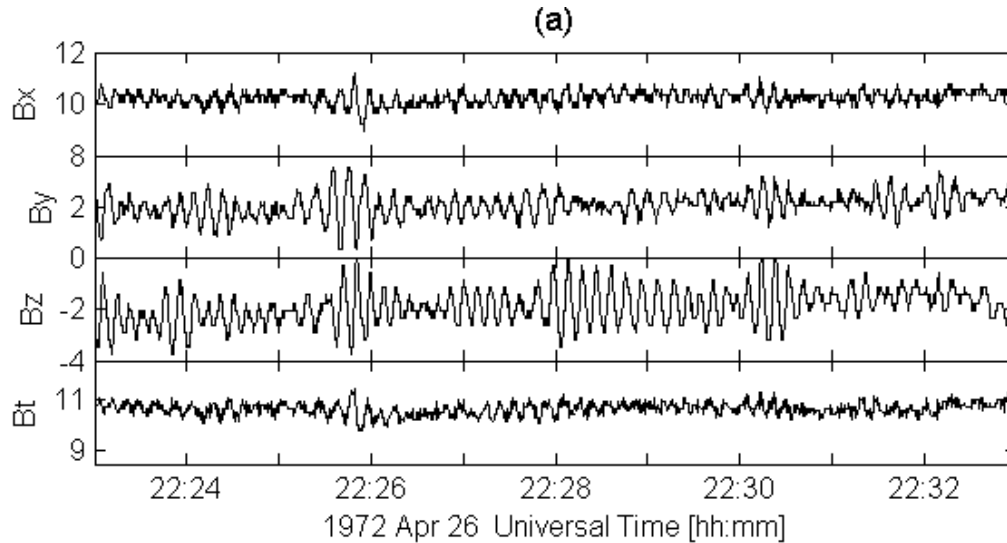
Apollo 12 LSM  
 Apollo 15 LSM  
 Apollo 16 LSM  
 Apollo 15 SBM  
 Apollo 16 SBM



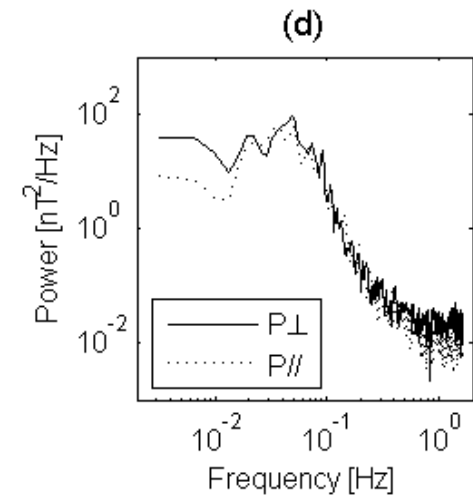
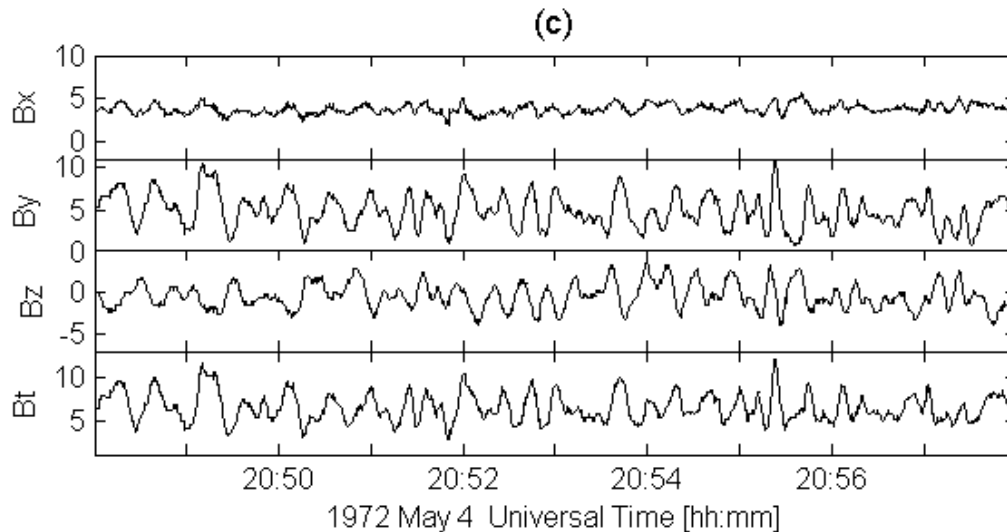
# Waves Observed by Apollo 15 LSM: Narrowband vs. Broadband

*Chi et al.*  
(2013), *PSS*.

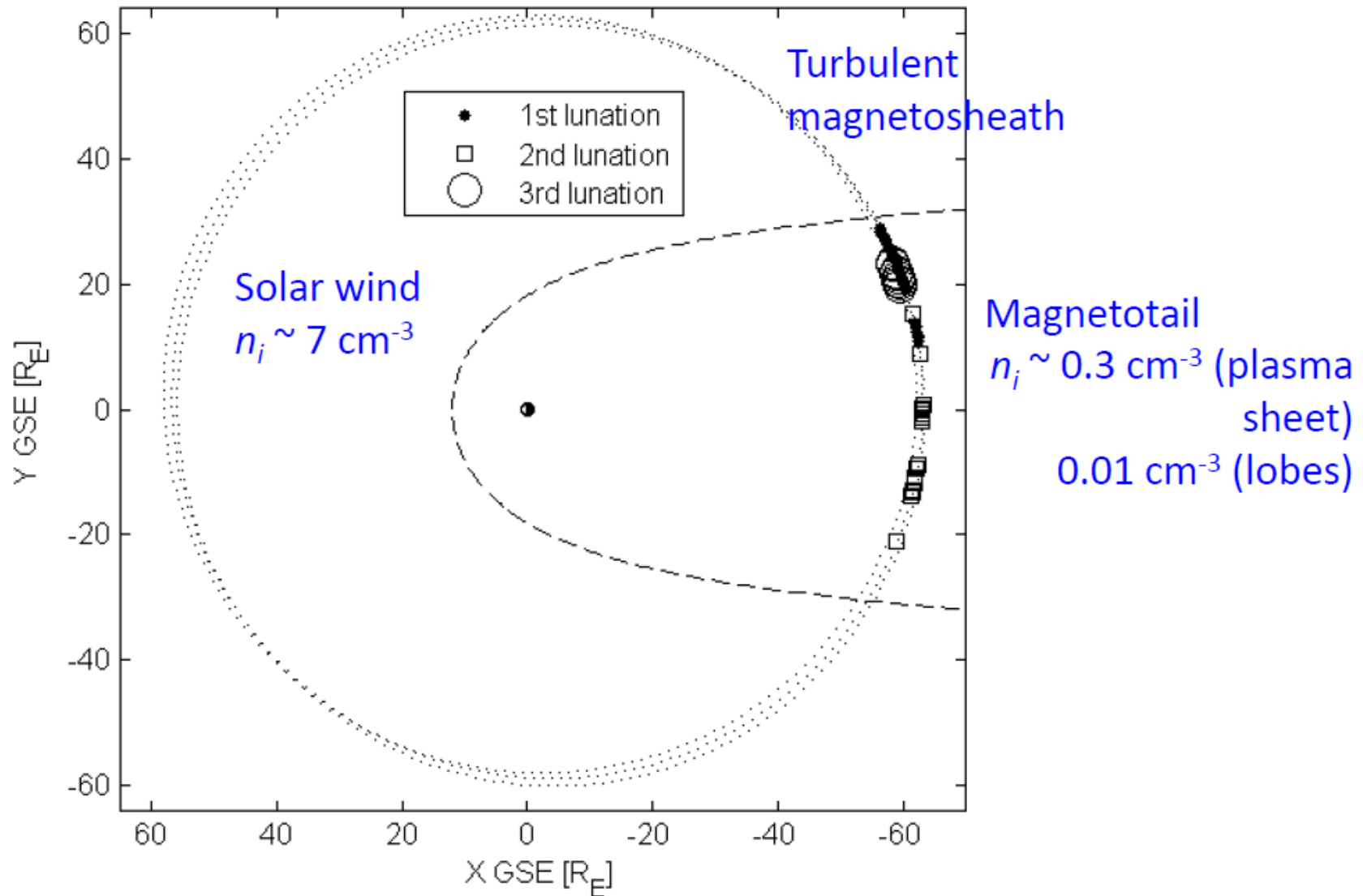
Magnetotail



Solar Wind

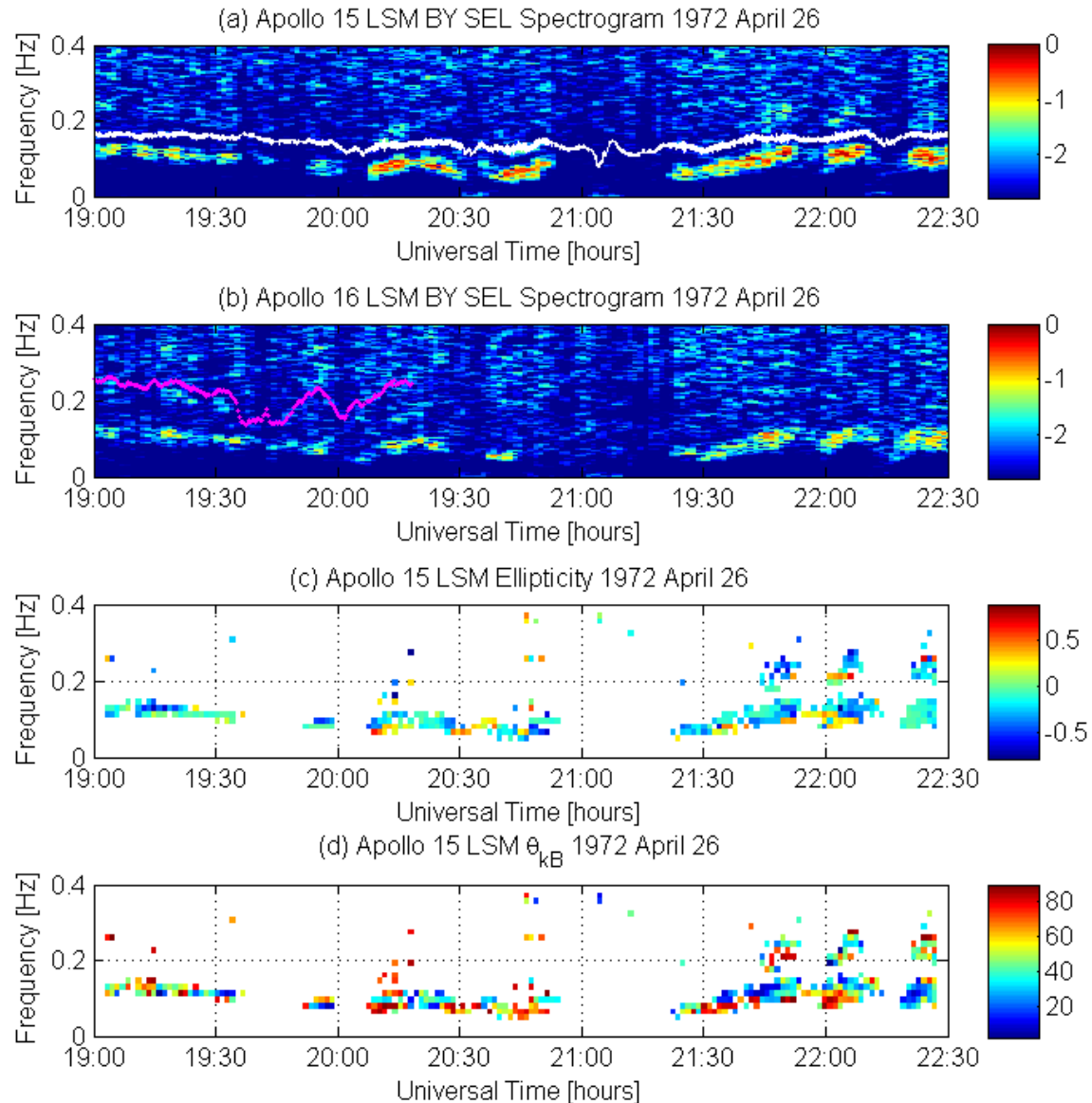


# Locations of Narrowband Waves



# Wave Amplitude and Polarization

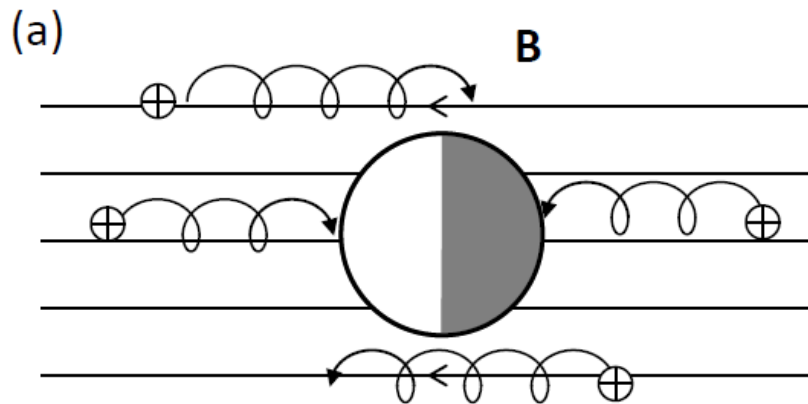
- Every narrow-band wave event was simultaneously observed at both Apollo sites.
- A16 LSM consistently observed lower wave amplitude (1/3 less).
- Second harmonics were sometimes seen at A15 LSM.
- Most waves are left-handed polarized.
- Propagation angles to B can be large.



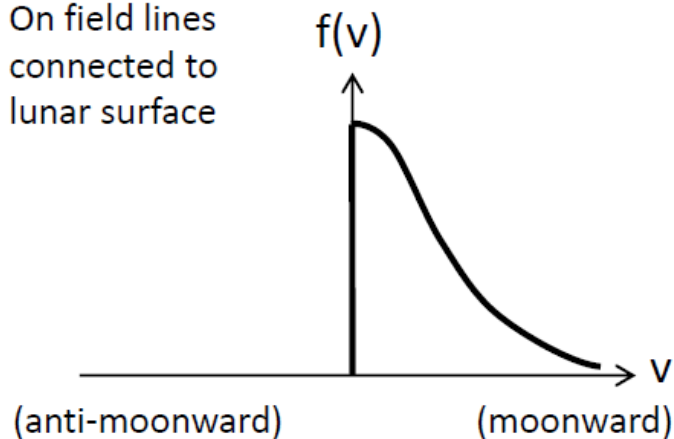
# Possible Mechanisms of Wave Excitation at the Moon

*Chi et al.  
(2013), PSS.*

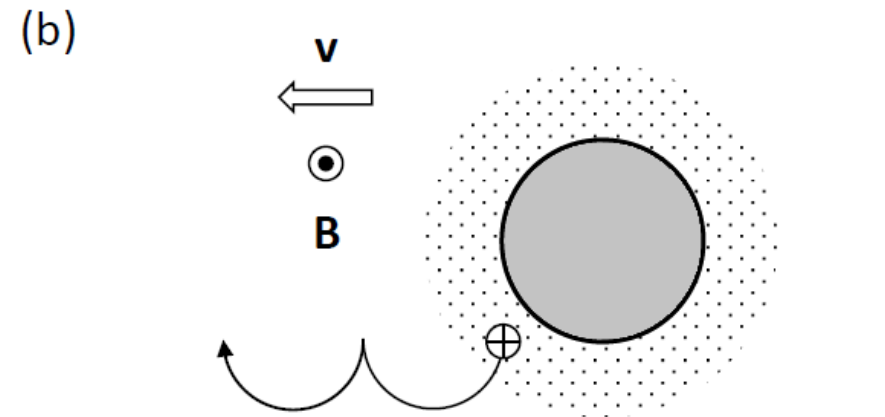
## *Ion Absorption at the Moon*



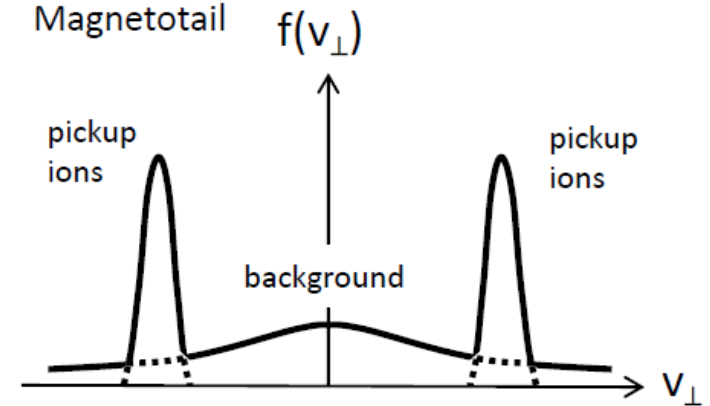
On field lines  
connected to  
lunar surface



## *Pickup Ions*



Magnetotail

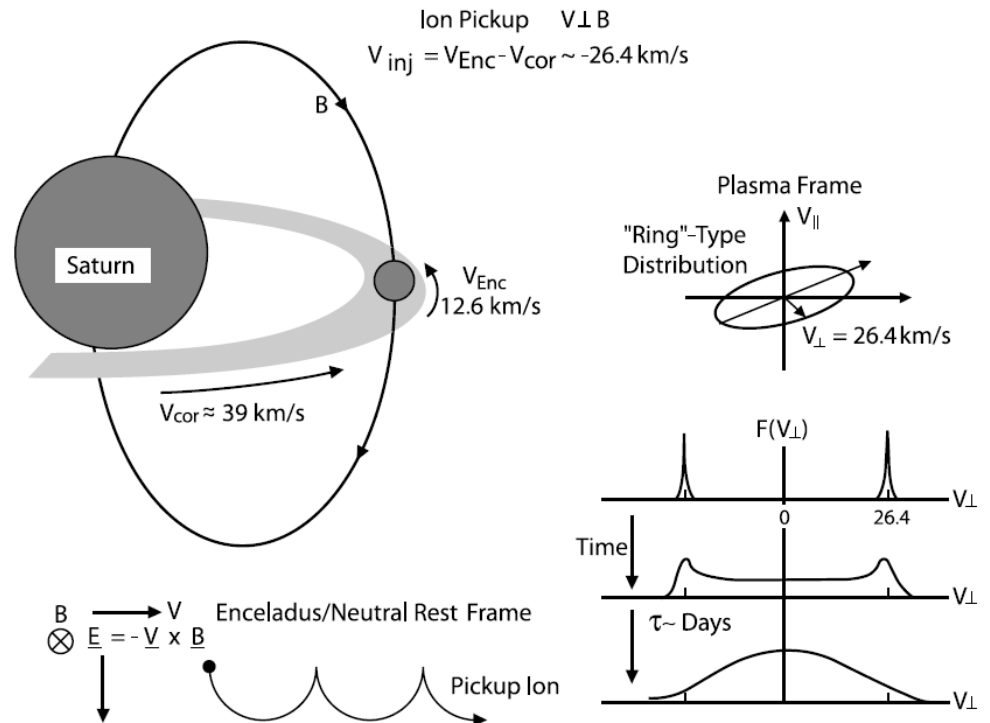


# PUI-excited Ion Cyclotron Waves in the Solar System

Ion cyclotron waves associated with pickup ions have been observed in the vicinity of:

- Mars exosphere
- Venus exosphere
- Neutral torus of Io
- E-ring of Saturn

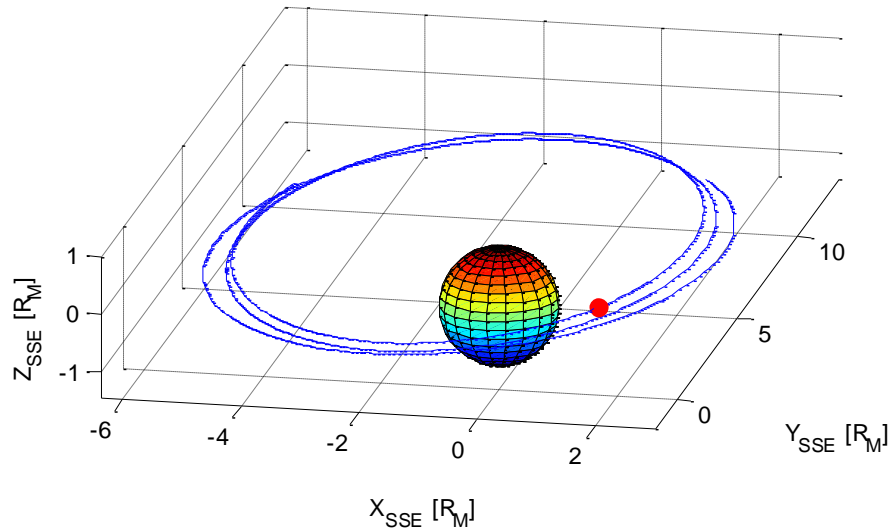
Observation of wave amplitude can be used to estimate the amount of the pick-up ions and atmospheric density.



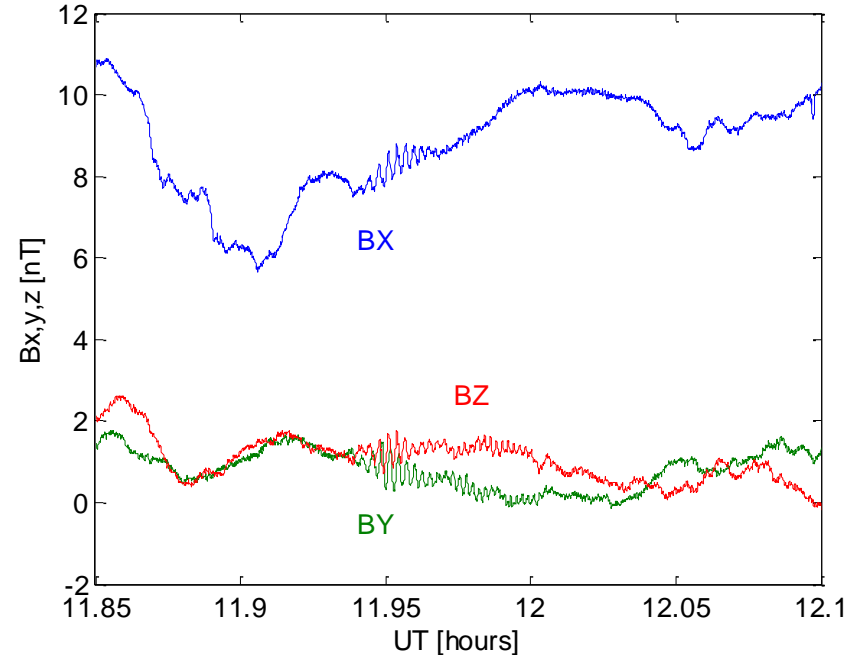
Cowee et al. [2009]

# ARTEMIS Observations of ICW: Event 1

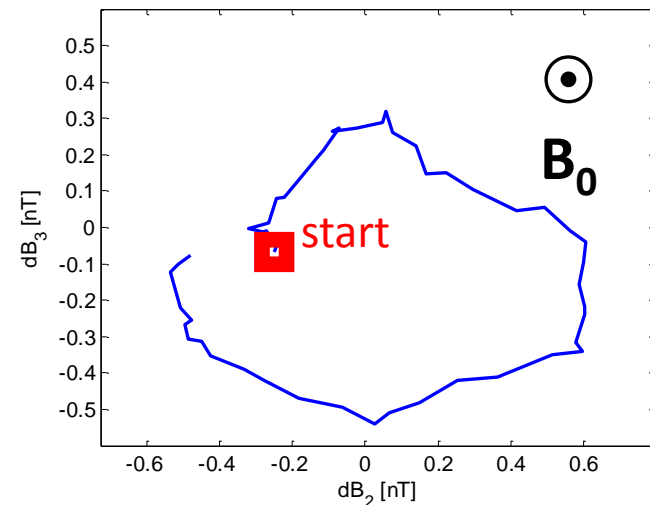
ARTEMIS TH-B 2011 Jul 13-16



ARTEMIS TH-B 2011-7-16

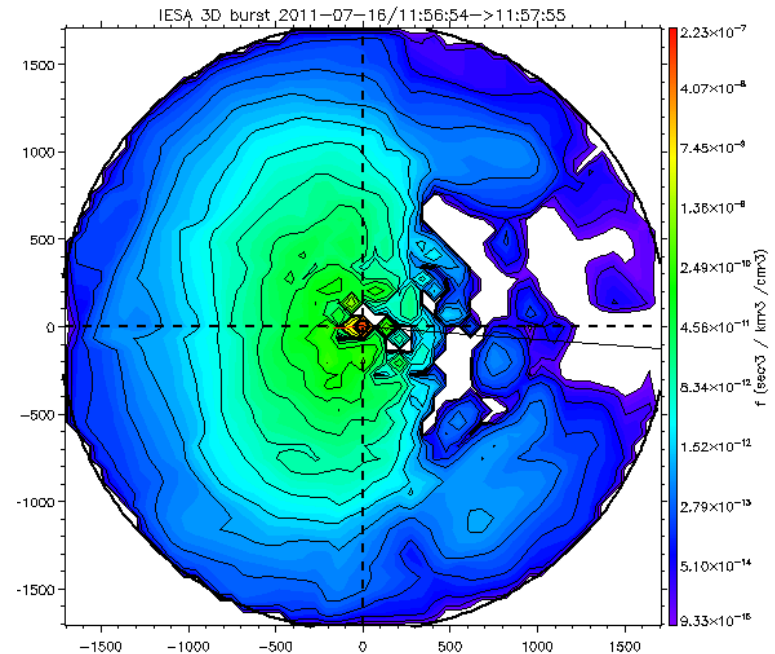
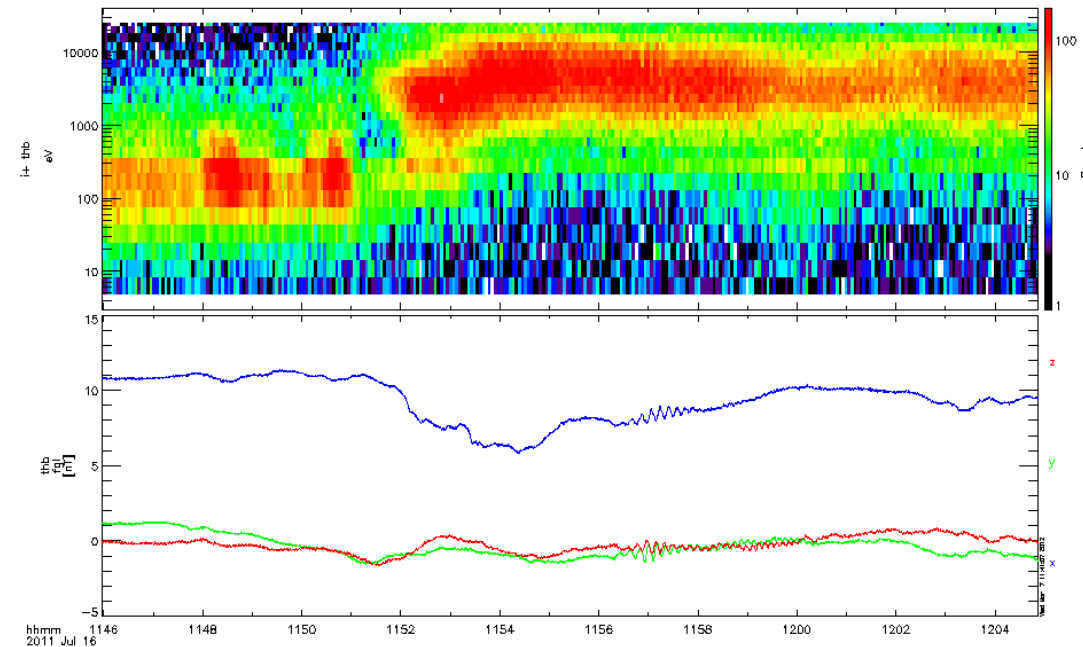


- Lunar spacecraft do not observe ion cyclotron waves as often.
- Several clear wave events were found close to the lunar dayside (highly location-dependent).
- $f \leq f_{cp}$
- Left-handed, elliptically polarized





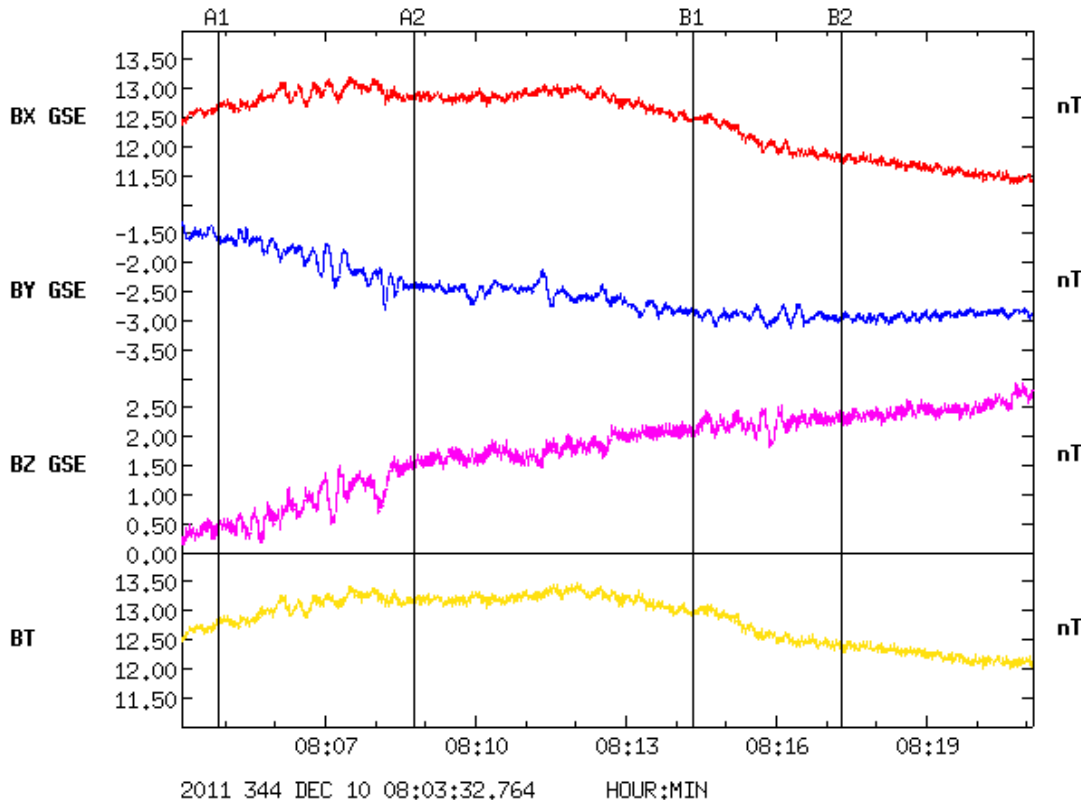
# Ion Energy/Velocity Distribution during Wave Event



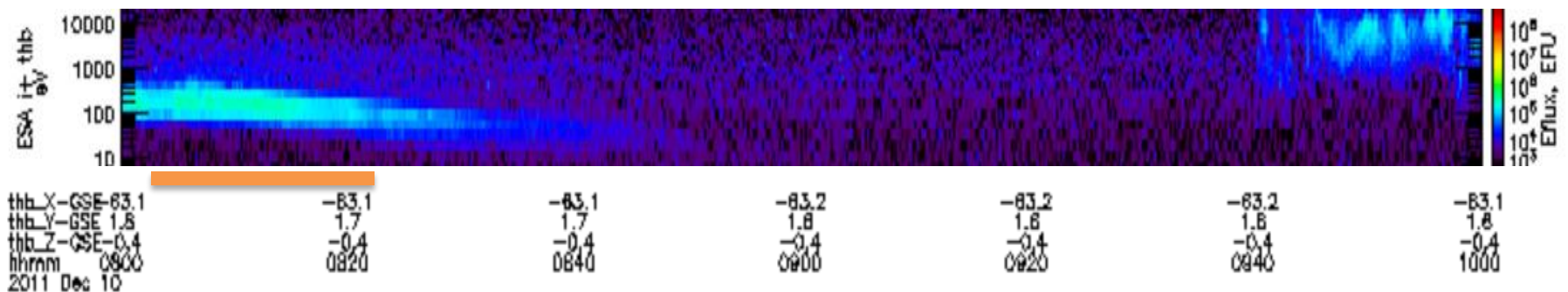
(from Jasper Halekas)

- Magnetic field line connected to the Moon
- Ion temperature  $\sim 3$  keV (plasma sheet)
- Ion cyclotron wave can be loss-cone driven

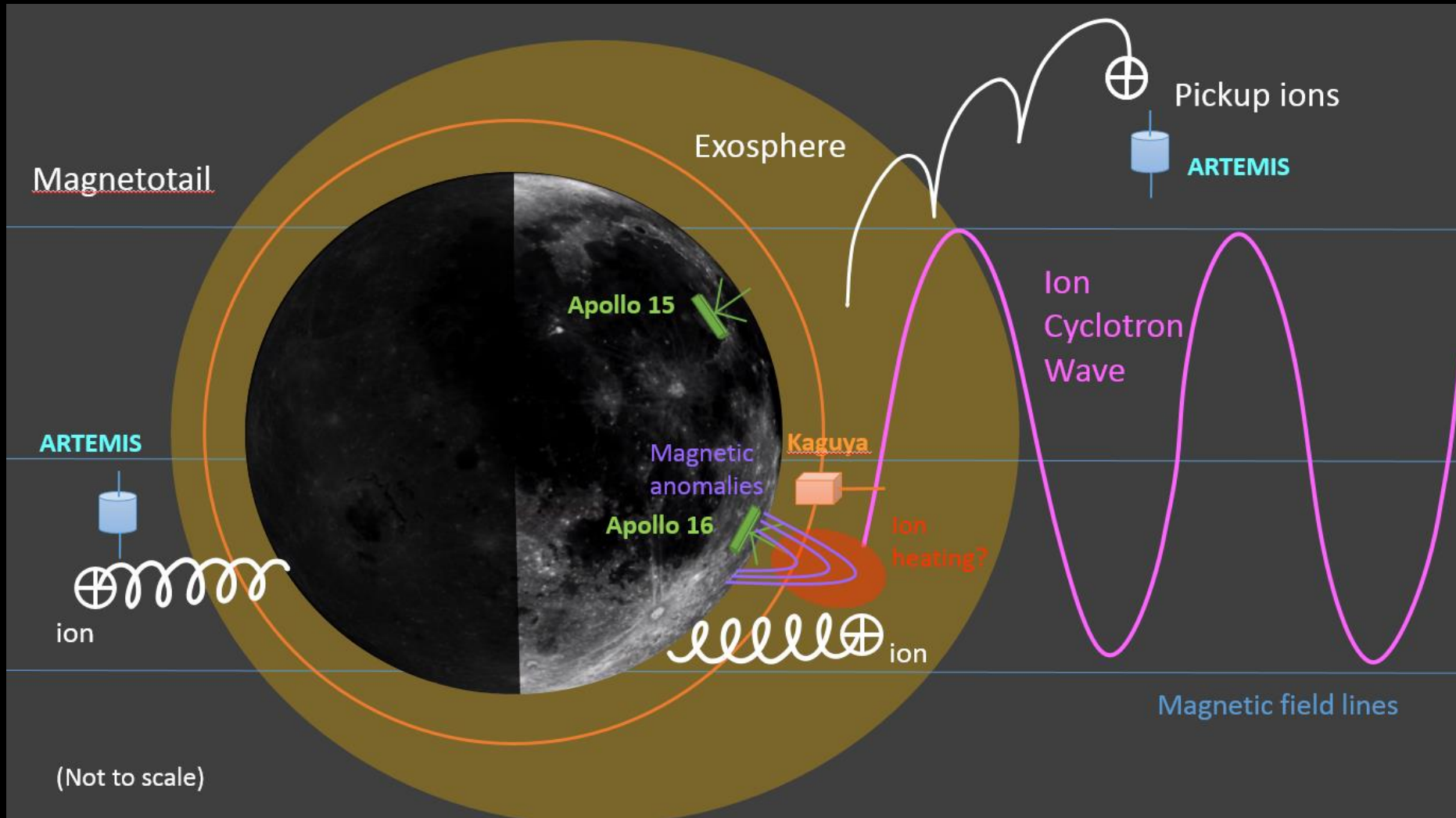
# ARTEMIS Observations of ICW: Event 2



- Past observations suggest He as a major species of lunar atmosphere.
- Observation was made by ARTEMIS in the magnetotail and 5.5 lunar radii from the Moon
- Wave properties:
  - $f$  at  $f_{c,He^+}$
  - Left-handed polarized
  - Field-aligned propagation
- Source particles are to be determined but pickup ions are one of the candidates.



# Magnetotail, Lunar Exosphere, ICW



# Main Points

- Interest: Understanding the two-way interaction between the Moon (the exosphere) and the magnetotail
- Relevance to FG: The tail environment at lunar distances can be influenced by the presence of the Moon. The Moon can be a dominant particle source in the tenuous magnetotail.
- Suggestions to Future Directions:
  - Identify the generation mechanism(s) of ion cyclotron waves at the Moon (through studying the morphology of ICW and the wave/particle data)
  - If the pickup ions are the source of ICW, the amount of exospheric particles (and their escape) can be estimated by the measurements of ICW.